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### Article

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# A remarkable recurrent nova in M31: Discovery and optical/UV observations of the predicted 2014 eruption (Corrigendum)

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An error has been identified in Table 3 of [Darnley et al. \(2015\)](#). The original table of fluxes included the incorrect unit ( $\times 10^{-15} \text{ W m}^{-2}$ ). The correct unit should have been ( $\times 10^{-15} \text{ erg cm}^{-2} \text{ s}^{-1}$ ), and hence a discrepancy of a factor of 1000 was introduced.

No other parts of the paper, nor the results reported within, were affected by this error. We have included a corrected version in Table 1.

## References

Darnley, M. J., Henze, M., Steele, I. A., et al. 2015, [A&A](#), 580, A45

**Table 1.** Selected observed emission lines and fluxes from the three epochs of Liverpool Telescope SPRAT spectra of the 2014 eruption of M31N 2008-12a.

Emission line	Flux <sup>a</sup> ( $\times 10^{-15} \text{ erg cm}^{-2} \text{ s}^{-1}$ )		
	$t = 0.32 \text{ d}$	$t = 1.44 \text{ d}$	$t = 2.17 \text{ d}$
H $\alpha$	$11.4 \pm 0.7$	$8.4 \pm 0.4$	$7.5 \pm 0.8$
H $\beta$	$3.1 \pm 0.1$	$2.2 \pm 0.2$	$0.6 \pm 0.3$
H $\gamma$	$2.0 \pm 0.4$	$1.5 \pm 0.2$	$0.5 \pm 1.0$
He I (7065 Å)	$3.0 \pm 0.4$	$1.9 \pm 0.3$	...
He I (6678 Å)	$2.0 \pm 0.4$	$1.3 \pm 0.4$	...
He I (5876 Å)	$2.5 \pm 0.3$	$1.7 \pm 0.2$	...

**Notes.** <sup>(a)</sup> Line flux is derived from the best-fit Gaussian profile for each emission line and is strongly dependent upon the adopted continuum level.