
Cloud structure of three Galactic infrared dark star-forming regions from combining ground and space based bolometric observations

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1. Introduction

Due to a mistake in proofreading, there is a sub-figure of Figure 11 missing in the paper, which corresponds to the second caption therein. The right figure and caption is below.

We also note that the reference empirical line of Molinari et al. (2008) in our Figure 7 is not accurately plotted, the slope of which should be \( \sim 1.13\)–1.27. While we do not use it to discuss our results since it is derived in a generally lower range of mass and luminosity of small-scale star-forming regions, we refer to the original work for a detailed description of this reference line.

Reference